

Tue 21 Mar

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Quantum statistics

↳ micro-states from energy levels occupation #s

Bosons $n_\ell = 0, 1, 2, \dots \rightarrow Z_g^{BE} = \prod_\ell \frac{1}{1 - e^{-\beta(E_\ell - \mu)}} \quad (\mu < E_\ell)$

Fermions $n_\ell = 0, 1$

Fermi-Dirac statistics for Fermions

Single energy level $E_0 \rightarrow Z_g = \sum_{n_0=0}^1 e^{-\beta(E_0 - \mu)n_0} = 1 + e^{-\beta(E_0 - \mu)}$

$E_i = E_0 n_0 \rightarrow \sum_\ell E_\ell n_\ell \quad N_i = \sum_\ell n_\ell$

General $E_\ell \rightarrow Z_g^{FD} = \sum_{n_0} \sum_{n_1} \dots \sum_{n_L} \exp\left[-\beta \sum_{\ell=0}^L (E_\ell - \mu)n_\ell\right]$

$= \left(\sum_{n_0} e^{-\beta(E_0 - \mu)n_0}\right) \dots \left(\sum_{n_L} e^{-\beta(E_L - \mu)n_L}\right)$ (factorization)

$= \prod_\ell (1 + e^{-\beta(E_\ell - \mu)})$ no need for $\mu < E$

product like $Z_N \sim Z_1^N$
but w/out over-counting

Collect results

$\Phi_{MB}(T, \mu) = -T \log \left(\prod_\ell \exp \left[e^{-\beta(E_\ell - \mu)} \right] \right) = -T \sum_\ell e^{-\beta(E_\ell - \mu)}$

$\Phi_{BE} = T \sum_\ell \log(1 - e^{-\beta(E_\ell - \mu)})$

$\Phi_{FD} = -T \sum_\ell \log(1 + e^{-\beta(E_\ell - \mu)})$

Classical (MB) physics should emerge from quantum statistics when # of accessible energy levels $\gg N$

Recall canonical spin system $\rightarrow p_i$ for $E_i > E_0$ exponentially suppressed for low T

$$p_i = \frac{1}{2} e^{-E_i/T} \rightarrow 0$$

\rightarrow Need high T for classical emergence

Let's compute $\langle N \rangle = \frac{-\partial \Phi}{\partial \mu}$ for all three & check high T

$$\begin{aligned} \langle N \rangle_{MB} &= T \sum_{\lambda} \frac{\partial}{\partial \mu} e^{-\beta(E_{\lambda} - \mu)} = \sum_{\lambda} \frac{1}{e^{\beta(E_{\lambda} - \mu)}} = \sum_{\lambda} \langle n_{\lambda} \rangle_{MB} \\ \langle N \rangle_{BE} &= -T \sum_{\lambda} \frac{-\beta e^{-\beta(E_{\lambda} - \mu)}}{1 - e^{-\beta(E_{\lambda} - \mu)}} = \sum_{\lambda} \frac{1}{e^{\beta(E_{\lambda} - \mu)} - 1} = \sum_{\lambda} \langle n_{\lambda} \rangle_{BE} \\ \langle N \rangle_{FD} &= T \sum_{\lambda} \frac{\beta e^{-\beta(E_{\lambda} - \mu)}}{1 + e^{-\beta(E_{\lambda} - \mu)}} = \sum_{\lambda} \frac{1}{e^{\beta(E_{\lambda} - \mu)} + 1} = \sum_{\lambda} \langle n_{\lambda} \rangle_{FD} \end{aligned}$$

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$$\langle n_{\lambda} \rangle_{MB} = \frac{1}{e^{\beta(E_{\lambda} - \mu)}}$$

$$\langle n_{\lambda} \rangle_{BE} = \frac{1}{e^{\beta(E_{\lambda} - \mu)} - 1}$$

$$\langle n_{\lambda} \rangle_{FD} = \frac{1}{e^{\beta(E_{\lambda} - \mu)} + 1} \leq 1 \checkmark$$

All occupation #s $\geq 0 \checkmark$

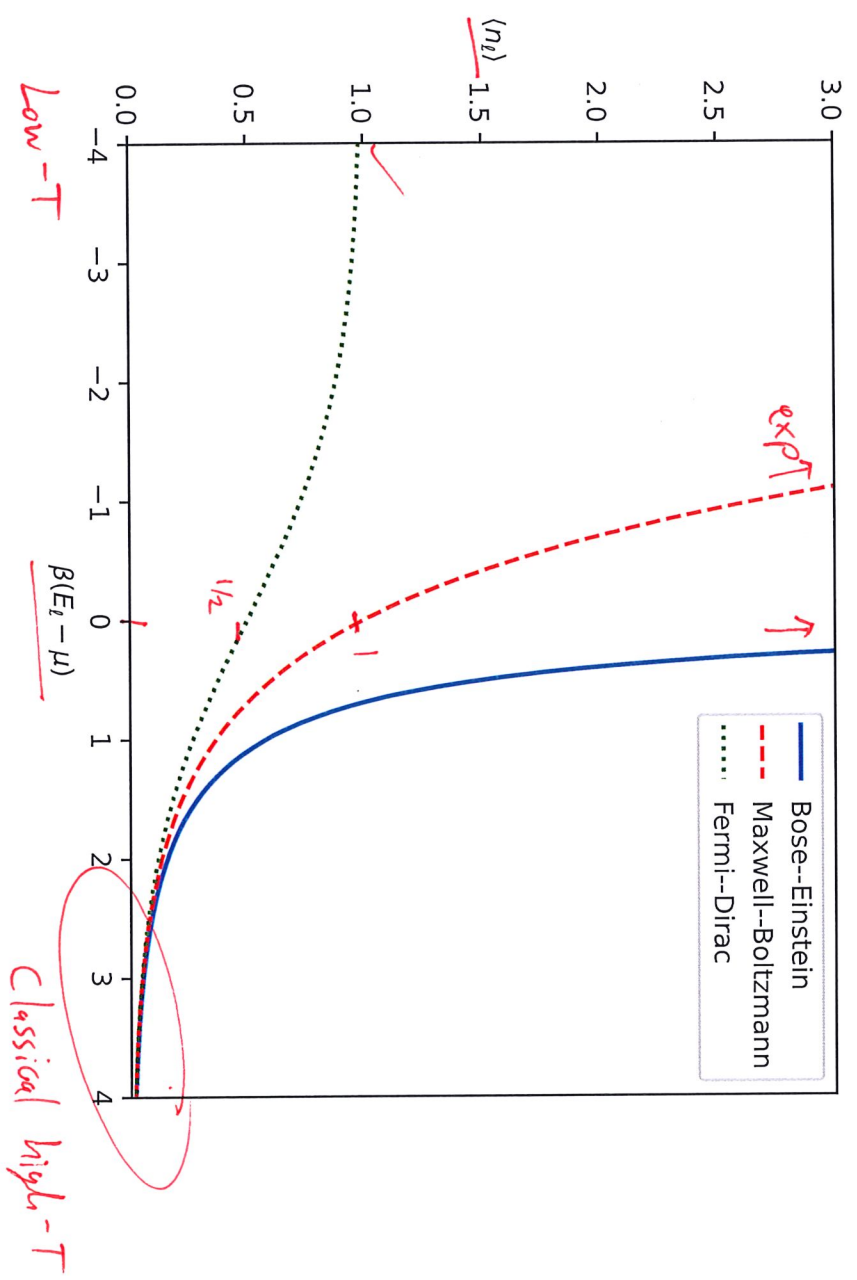
All agree when $e^{\beta(E_{\lambda} - \mu)} \gg 1 \rightarrow \langle n_{\lambda} \rangle \ll 1 \checkmark$

Classical limit $\beta(E_{\lambda} - \mu) \gg 1$ high T ???

$\beta \rightarrow 0$ w/fix $(E_{\lambda} - \mu)$

$$\langle n_{\lambda} \rangle_{MB} \approx 1 \quad \langle n_{\lambda} \rangle_{FD} \approx \frac{1}{2} \quad \langle n_{\lambda} \rangle_{BE} \rightarrow \infty$$

True classical limit is $\{T, -\mu\} \rightarrow \infty$ s.t. $-\mu \gg T \gg E_{\lambda}$
 (Fewer particles, more accessible energy levels)



Look ahead to quantum gases — $\left\{ \begin{array}{l} \text{grand-canonical ensemble} \\ \text{quantum statistics} \end{array} \right.$

Need energies E_e going into Z_g^{BE} and Z_g^{FD}

Two limiting cases

non-relativistic

$$E = \frac{p^2}{2m}$$

$$= \frac{\hbar^2 \pi^2}{2mL^2} (k_x^2 + k_y^2 + k_z^2)$$

$$k_{x,y,z} = 1, 2, 3, \dots$$

Heisenberg uncertainty

$$(\Delta x)(\Delta p_x) \gtrsim \hbar$$

$$\Delta x \leq L$$

$$\Delta p_x \gtrsim \frac{\hbar}{L}$$

ultra-rel.

$$m \rightarrow 0$$

$$E = cp = \hbar c \frac{\pi}{L} k$$

$$\sqrt{p_x^2 + p_y^2 + p_z^2}$$

Waves \rightarrow k is wave number

$$\frac{2L}{k} = \lambda \text{ is quantized wavelength}$$

$$\omega = 2\pi f$$

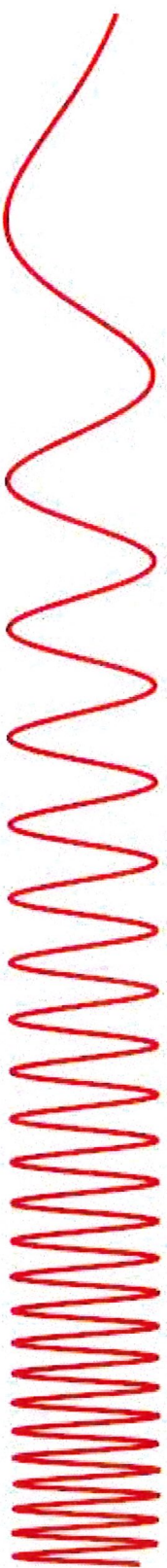
$$= \frac{2\pi}{\lambda} c = c \frac{\pi}{L} k$$

$$E = \hbar \omega$$

Large $E \rightarrow$ high freq
short λ
"ultraviolet"

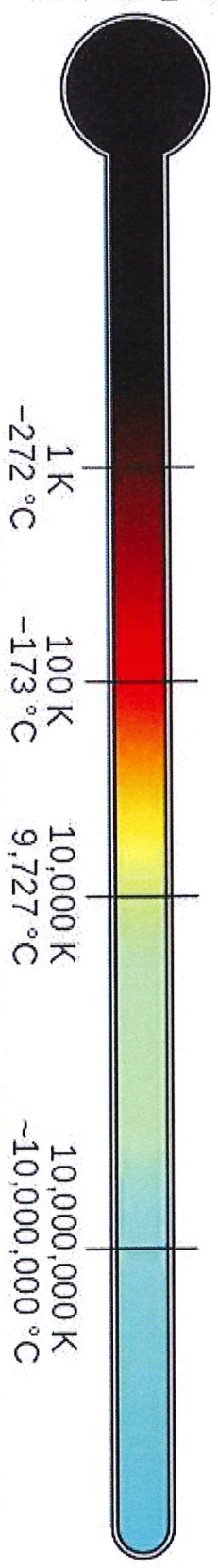
Small $E \rightarrow$ low freq.
long λ
"infrared"

Penetrates Earth's Atmosphere?



Radiation Type	Approximate Scale of Wavelength	Frequency (Hz)	Examples
Radio	10^3 m	10^4 Hz	Buildings
Microwave	10^{-2} m	10^8 Hz	Humans
Infrared	10^{-5} m	10^{12} Hz	Butterflies, Needle Point
Visible	0.5×10^{-6} m	10^{15} Hz	Protozoans
Ultraviolet	10^{-8} m	10^{16} Hz	Molecules
X-ray	10^{-10} m	10^{18} Hz	Atoms
Gamma ray	10^{-12} m	10^{20} Hz	Atomic Nuclei

Temperature of objects at which this radiation is the most intense wavelength emitted



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