

MATH327: Statistical Physics

Monday, 27 February 2023

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Something to consider

Last week we met the canonical partition function,

$$Z = \sum_i e^{-E_i/T},$$

which sums over all micro-states ω_i .

What can we do if there are uncountably many micro-states?

Recap

Micro-canonical temperature behaves sensibly

Canonical ensemble

Based on partition function $Z = \sum_i e^{-\beta E_i}$

or Helmholtz free energy $F = -T \log Z$

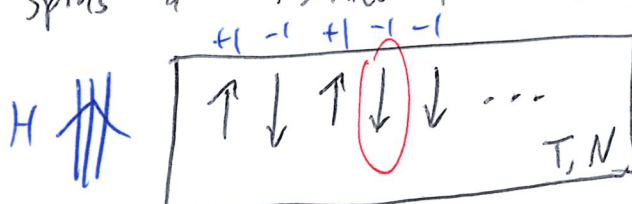
derived energy $\langle E \rangle$, entropy S , heat capacity c_v

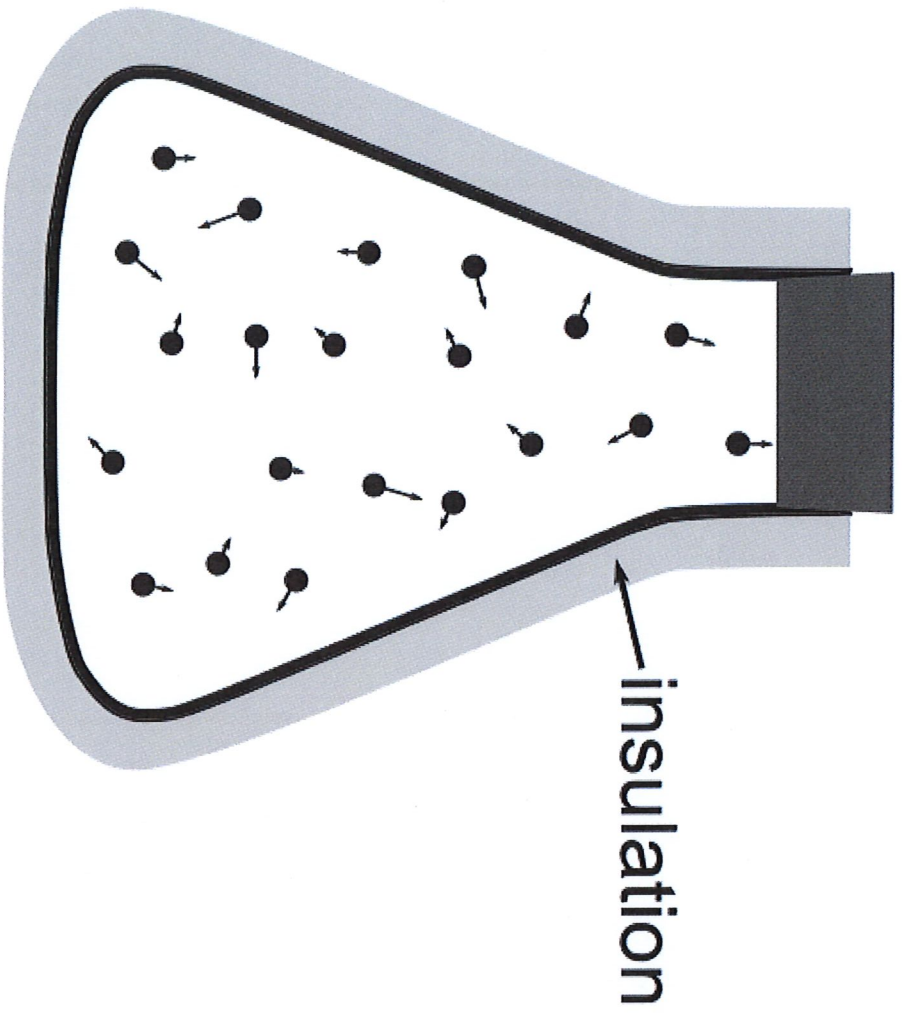
Application: Information

Pure info content \rightarrow physically observable effects
knowable in principle

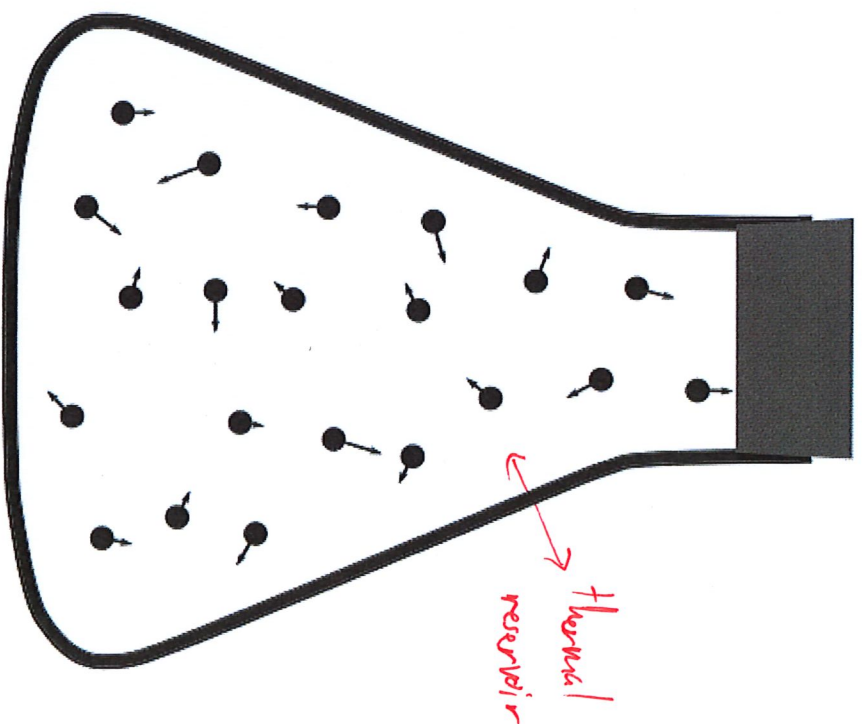
Simple example: Spin systems with $\left\{ \begin{array}{l} \text{distinguishable} \\ \text{indistinguishable} \end{array} \right\}$ spins

Distinguishable spins at fixed locations in solid





Microcanonical
(const. N E)



Canonical
(const. N T)

$M = 2^N$ micro-states w_i with energies E_i
and probabilities $p_i = \frac{1}{Z} e^{-E_i/T}$

Call parallel spin $s_n = 1$
anti-parallel spin $s_n = -1$

Then $E_i = -H \sum_{n=1}^N s_n$ for w_i specified by $N \{s\}$

Start with the partition function

$$Z_D = \sum_i e^{-\beta E_i} = \sum_{s_1=\pm 1} \dots \sum_{s_N=\pm 1} \exp\left[\beta H \sum_n s_n\right] \quad x = \beta H = \frac{H}{T}$$

$$= \left(\sum_{s_1=\pm 1} e^{x s_1} \right) \dots \left(\sum_{s_N=\pm 1} e^{x s_N} \right)$$

$$= \left(\sum_{s=\pm 1} e^{x s} \right)^N = (e^x + e^{-x})^N = [2 \cosh(\beta H)]^N$$

(Factorization)

Helmholtz Free energy $F_D(\beta) = -\frac{\log Z_D}{\beta} = \frac{-N \log [2 \cosh(\beta H)]}{\beta}$

$$\langle E \rangle_D = -T^2 \frac{\partial}{\partial T} \left(\frac{F}{T} \right) = \frac{\partial}{\partial \beta} (\beta F)$$

$$= -N \frac{\partial}{\partial \beta} \log [2 \cosh(\beta H)] = \frac{-N}{2 \cosh(\beta H)} (2 \sinh(\beta H)) \cdot H$$

$$= -NH \tanh(\beta H)$$

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Entropy $S_D = \beta (\langle E \rangle_D - F_D) = -NH \tanh(\beta H) + N \log [2 \cosh(\beta H)]$
 $\log(e^{\beta H})$

Low-temperature limit

$$\beta \rightarrow \infty \quad \langle E \rangle_D = -NH \tanh(\beta H) \rightarrow -NH = E_0$$

(ground state) ✓

Absolute zero temperature \rightarrow single ground state

Zero entropy

$$S_D \rightarrow -NH + NH = 0$$

Low-temperature expansions

$T > 0$ allows contributions from higher-energy micro-states
 $E_i > E_0$ "excited states"

suppressed $\sim P_i = e^{-E_i/T}$

Spin system \rightarrow energy levels separated by constant energy gap

$$\Delta E = E_{n+1} - E_n = 2H$$

$$E = -H(2n+1)$$

Gap controls approach to zero-temperature limits

$$\begin{aligned} \frac{\langle E \rangle_0}{NH} &= -\tanh(\beta H) \quad \text{expand } \beta H \gg 1 \quad e^{-2\beta H} \ll 1 \\ &= -\frac{1 - e^{-2\beta H}}{1 + e^{-2\beta H}} = -(1 - e^{-2\beta H}) (1 - e^{-2\beta H} + O(e^{-4\beta H})) \\ &= -1 + 2e^{-2\beta H} + O(e^{-4\beta H}) \\ &= -1 + 2e^{-\Delta E/T} + O(e^{-2\Delta E/T}) \end{aligned}$$

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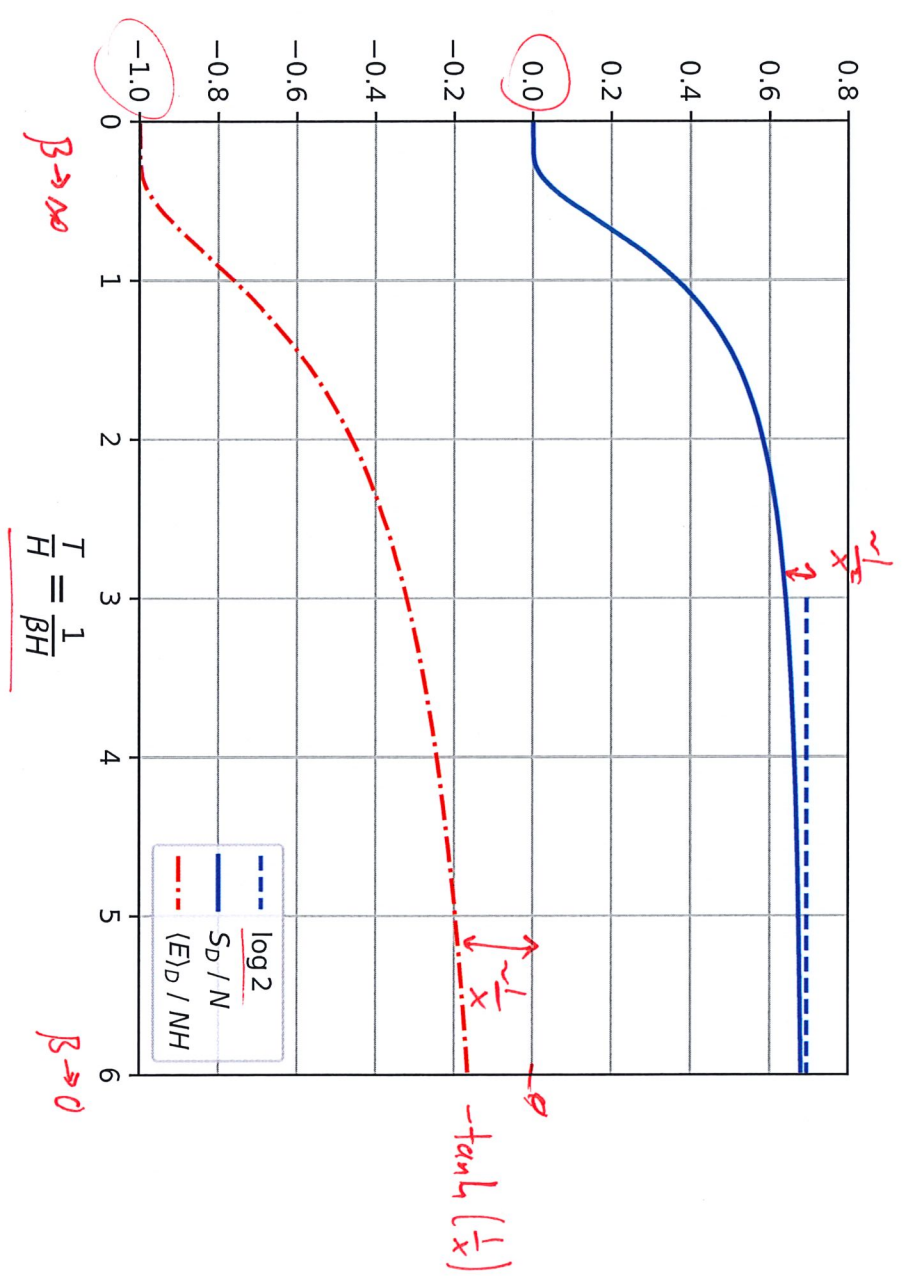
(exponential suppression of excited-state effects)

$$\frac{P_{n+1}}{P_n} = \frac{e^{-(E_{n+1})/T}}{e^{-E_n/T}} = e^{-\beta \Delta E}$$

Similarly

$$\begin{aligned} \frac{S_D}{N} &= -\beta H \tanh(\beta H) + \log [2 \cosh(\beta H)] \\ &= -\beta H \frac{1 - e^{-2\beta H}}{1 + e^{-2\beta H}} + \log [2 \cosh(\beta H)] \\ \log [2 \cosh(\beta H)] &= \log [e^{\beta H} (1 + e^{-2\beta H})] = \beta H + \log(1 + e^{-2\beta H}) \\ &= \beta H + e^{-2\beta H} + O(e^{-4\beta H}) \\ \frac{S_D}{N} &= \beta \Delta E \frac{1 - e^{-\beta \Delta E}}{1 + e^{-\beta \Delta E}} + e^{-\beta \Delta E} + O(e^{-2\beta \Delta E}) \end{aligned}$$

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High-temperature expansions

$$\frac{\langle E \rangle_0}{NH} = -\tanh(\beta H) = -\beta H + \frac{(\beta H)^3}{3} + \mathcal{O}[(\beta H)^5] \rightarrow 0 \quad \text{as } \beta \rightarrow 0 \quad (T \rightarrow \infty)$$

$\beta H \ll 1$

(Recall micro-canonical $T \rightarrow \infty$ for conserved $E \rightarrow 0$)

$$\frac{S_D}{N} = -\beta H \tanh(\beta H) + \log [2 \cosh(\beta H)]$$

$$= \log 2 + \log \left(1 + \frac{1}{2}(\beta H)^2 + \mathcal{O}[(\beta H)^4] \right)$$

$$= \log 2 + \frac{1}{2}(\beta H)^2 + \mathcal{O}[(\beta H)^4]$$

$$\frac{S_D}{N} = \log 2 - \frac{1}{2}(\beta H)^2 + \mathcal{O}[(\beta H)^4]$$

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Note $S_D \rightarrow N \log 2 = \log 2^N = \log M$ (micro-canonical)

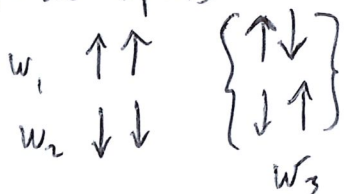
as $T \rightarrow \infty$

$$\text{all } P_i = \frac{1}{Z} e^{-E_i/T} \approx \frac{1}{Z} \approx \frac{1}{M} \quad Z = \sum_i e^{-E_i/T} \approx M$$

Indistinguishable spins in freely moving gas (one-dim.)



$N=2$ spins \rightarrow three micro-states, not 2^N



In general, can only know total $\{n_+, n_-\}$ for each w_i

One micro-state for each energy level!

Example: For $N=4$, the micro-states are $E = -4H, -2H, 0, 2H, 4H$
all up all down

$N+1$ micro-states in general

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Label micro-state w_k with $E_k = -NH + 2Hk = -H(N-2k)$
 ($\langle E \rangle$) $k=0, 1, \dots, N$

Start with partition Function

$$Z_I = \sum_{k=0}^N e^{-E_k/T} = \sum_{k=0}^N e^{\beta H(N-2k)} = e^{N\beta H} \sum_{k=0}^N (e^{-2\beta H})^k$$

$$\sum_{k=0}^N x^k = \sum_{k=0}^{\infty} x^k - \sum_{k=N+1}^{\infty} x^k = \frac{1}{1-x} - x^{N+1} \sum_{l=0}^{\infty} x^l$$

$$= \frac{1 - x^{N+1}}{1-x}$$

$$x = e^{-2\beta H} < 1$$

$$Z_I = e^{N\beta H} \left(\frac{1 - e^{-2(N+1)\beta H}}{1 - e^{-2\beta H}} \right)$$

~~Helmholtz~~ Helmholtz Free energy

$$F_I = \frac{-\log Z_I}{\beta} = -NH - \frac{1}{\beta} \log(1 - e^{-2(N+1)\beta H}) + \frac{1}{\beta} \log(1 - e^{-2\beta H})$$

Clearly different vs. $F_D = -\frac{N}{\beta} \log[2 \cosh(\beta H)]$

↳ Lead to different $\langle E \rangle_I$ and S_I (homework)

Same $T \rightarrow 0$ limits - single ground state

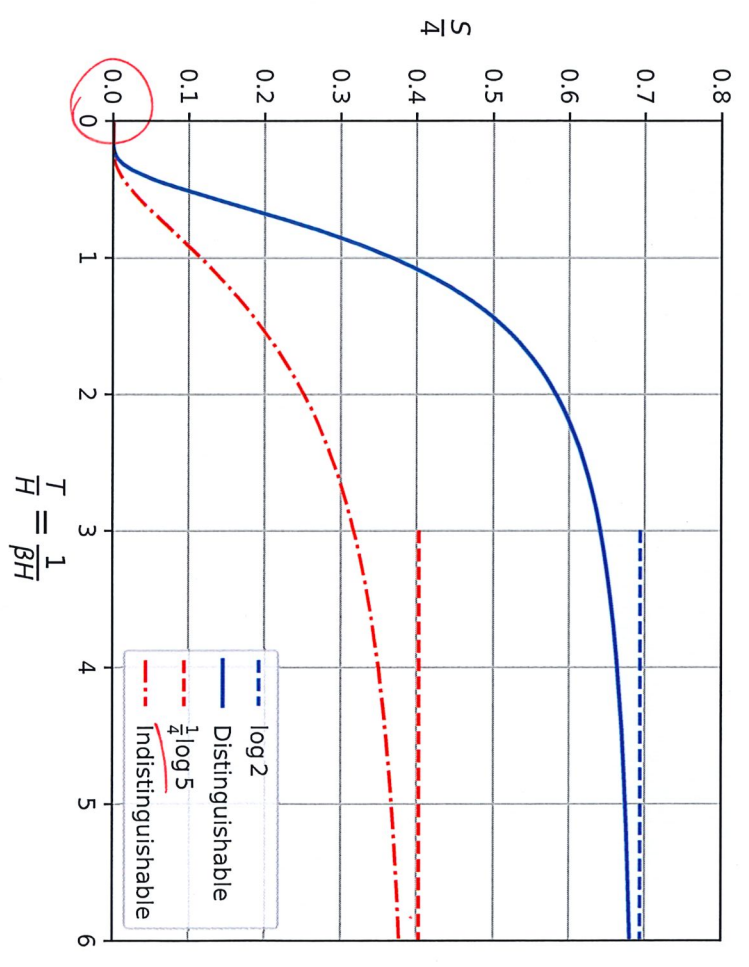
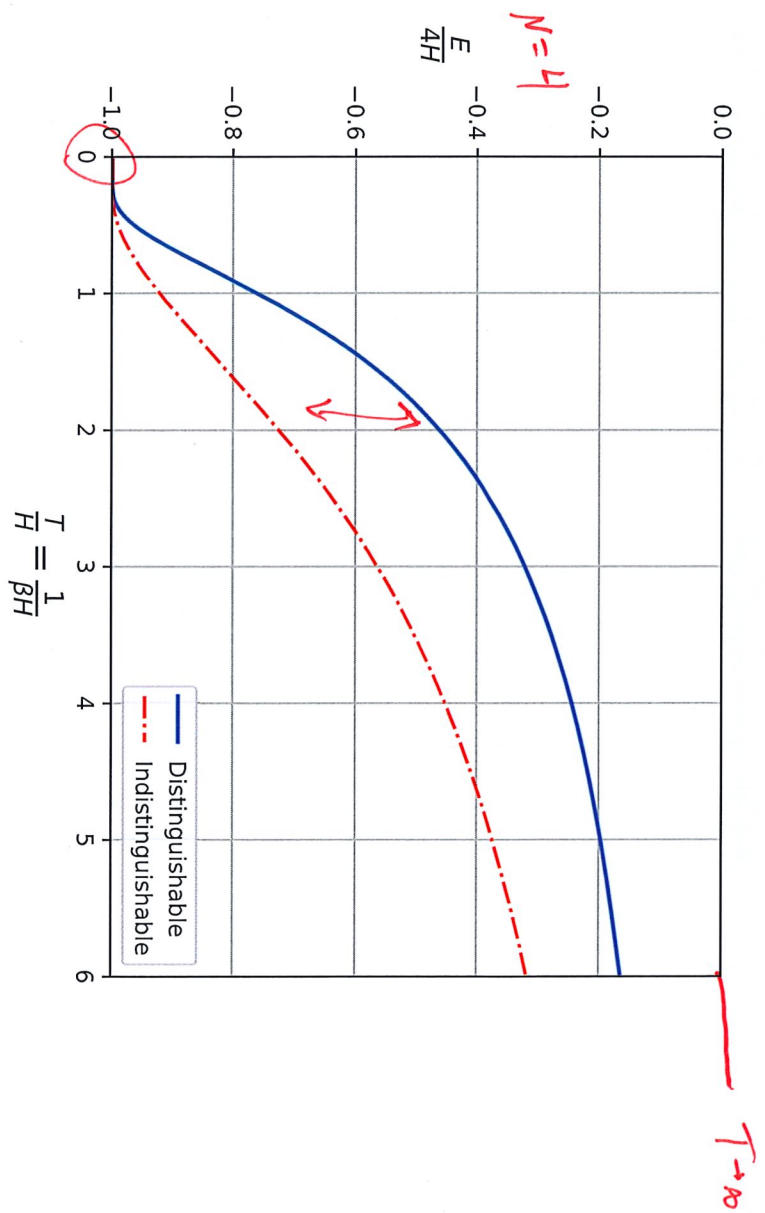
Energy $\rightarrow 0$ as $T \rightarrow \infty$

High-temperature $S_I \rightarrow \log(N+1) = \log M$

exponentially fewer micro-states
 $N+1$ vs. $2^N = \exp(N \log 2)$

Physically measurable effects from intrinsic info content

"Information \Rightarrow is physical"



Application: Ideal gases

Classical, non-relativistic, ideal gases

↳ Not quantum, exactly know (x, y, z) and $\vec{p} = (p_x, p_y, p_z)$

Non-rel. means slow particles (compared speed of light)

$$E_n = \frac{1}{2m} P_n^2 \quad \text{For mass } m$$

inner product $P_n^2 = \vec{p}_n \cdot \vec{p}_n = p_x^2 + p_y^2 + p_z^2$

Ideal means no interactions between particles

$$E = \frac{1}{2m} \sum_{n=1}^N P_n^2$$

Put N particles in cubic box of volume $V = L^3$

Temperature T fixed by thermal reservoir

Start with partition function

$$Z = \sum_i e^{-E_i/T}$$

problem

E_i from continuous \vec{p}_n

Need to regulate system

↳ countable micro-states with well-defined partition func

Then limit of sum \rightarrow integral

Declare that only possible momenta are

$$\vec{p} = (p_x, p_y, p_z) = \hbar \frac{\pi}{L} (k_x, k_y, k_z)$$

$$k_{x,y,z} = 0, 1, 2, \dots$$

countable

Planck constant

unit conversion from $\frac{1}{L}$ to p

Similar "quantized" momenta realized of nature

Here just ansatz

Energies are also countable

$$E_k = \frac{p^2}{2m} = \frac{\hbar^2 \pi^2}{2mL^2} (k_x^2 + k_y^2 + k_z^2) \Rightarrow E = \frac{\hbar^2 \pi^2}{2mL^2}$$

Lowest energies

$$E = E_k \{0, 1, 2, 3, 4, 5, \underline{6}, 8, 9, 10, 11, 12, 13, 14, \dots\}$$

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No 7, 15, ~~16~~ ...