

New strong dynamics beyond the standard model

Lecture 6

5 December 2017

Last time

- Lattice discretization of spacetime regularizes strongly coupled gauge theories
Lattice spacing $a \rightarrow$ UV cutoff $\Lambda_{UV} = 1/a$, removed in continuum limit $a \rightarrow 0$
- Finite lattice \rightarrow observables from numerically evaluating functional integral
- Must know fundamental lagrangian at cutoff scale a^{-1} to formulate lattice theory
- Gauge fields \rightarrow gauge links $U \sim e^{-igaA}$, closed loops of links are gauge invariant
- Lattice fermion discretizations either break chiral symmetry or produce ‘doubblers’
 - Naive:** $16F$ continuum fermions, $U(4F)_V \times U(4F)_A$ chiral symmetry
 - Staggered:** $4F$ continuum fermions, $U(F)_V \times U(F)_A$ chiral symmetry
 - Wilson:** F continuum fermions, no chiral symmetry (explicitly broken)
 - Domain wall:** F fermions, preserves lattice ‘remnant’ of $SU(F)_V \times SU(F)_A$
- Two-point correlators $C(t) \rightarrow$ composite particle masses and decay constants
Representative Wilson and staggered results for some systems beyond QCD

S parameter from lattice vacuum polarization $\Pi_{V-A}(Q^2)$

- In lecture 2 we saw $S = 4\pi [\Pi'_{VV}(0) - \Pi'_{AA}(0)]_{\text{new}} = 4\pi \lim_{Q^2 \rightarrow 0} \frac{d}{dQ^2} \Pi_{V-A}^{(\text{new})}(Q^2)$

Vacuum polarization tensor $\Pi^{\mu\nu}(x)$ is also two-point correlation function,

but now we Fourier transform rather than projecting to $\vec{p} = 0$

$$\Pi_{V-A}^{\mu\nu}(Q) = Z \sum_n e^{iQ \cdot (n + \hat{\mu}/2)} \text{Tr} [\langle \mathcal{V}_\mu(n) V_\nu(0) \rangle - \langle \mathcal{A}_\mu(n) A_\nu(0) \rangle]$$

$$\Pi^{\mu\nu}(Q) = \left(\delta^{\mu\nu} - \frac{Q^\mu Q^\nu}{Q^2} \right) \Pi(Q^2) - \frac{Q^\mu Q^\nu}{Q^2} \Pi^L(Q^2)$$

Smallest accessible $Q^2 = \left(\frac{2\pi n}{N_t} \right)^2, \left(\frac{2\pi n}{N_s} \right)^2$ with $n = 1, 2, \dots$

- Need domain wall fermions to ensure same renormalization factor $Z_V = Z_A \equiv Z$
- **DWF conserved current** \mathcal{V} sums over N_5 (similarly for \mathcal{A} , with $\text{sign}[s - N_5/2]$),

$$\mathcal{V}_\mu(x) = \sum_{s=1}^{N_5} \left[\bar{\Psi}(x + \hat{\mu}, s) \frac{1 + \gamma_\mu}{2} U_\mu^\dagger(x) \Psi(x, s) - \bar{\Psi}(x, s) \frac{1 - \gamma_\mu}{2} U_\mu(x) \Psi(x + \hat{\mu}, s) \right]$$

Ward identity would give $Z = 1$, but $V_\mu(x) = \bar{\psi}(x) \gamma_\mu \psi(0)$ on walls $\rightarrow Z \neq 1$

- Non-perturbatively determine Z by comparing correlators involving \mathcal{A}_μ vs. A_μ

S parameter from lattice $\Pi_{V-A}(Q^2)$ (continued)

- **Extract slope** $\Pi'_{V-A}(Q^2 = 0)$ by fitting to generalized single pole dominance

$$\Pi_{V-A}(Q^2) \simeq -F_P^2 + \frac{Q^2 F_V^2}{Q^2 + M_V^2} - \frac{Q^2 F_A^2}{Q^2 + M_A^2} \longrightarrow \frac{a_0 + a_1 Q^2}{1 + b_1 Q^2 + b_2 Q^4} \quad (1)$$

(Weinberg sum rule $F_V^2 - F_A^2 = F_P^2$ motivates dropping Q^4 term in numerator)

- **Result:** S decreases compared to QCD as N_F increases,

if extra PNCBs have $M_P < M_{V0}$ but small chiral logs $\propto (N_F - 2) \log\left(\frac{M_{V0}^2}{M_P^2}\right)$

No symmetry breaking for sufficiently large $N_F \gtrsim 10 \longrightarrow$ smaller S reasonable

In practice, harder to reach light masses $M_P < M_{V0}$ as N_F increases

- **Future:** EW gauge boson contribution to Higgs potential, $c_V \propto \int dQ^2 \Pi_{U-B}(Q^2)$
Fermion contributions **involve four-point correlators** of baryonic operators

Supplement: Lattice calculations of LECs

- In lecture 2 we saw the S parameter is an EFT low-energy coefficient (LEC)

It corresponds to ℓ_5 in the SU(2) chiral lagrangian,

equivalently α_1 in the electroweak chiral lagrangian (without light Higgs)

- Computations of EFT low-energy coefficients are common lattice projects
(motivating domain wall fermions with continuum-like global symmetries)

- **Example:** $\langle \bar{\psi}\psi \rangle / F^3 \propto B/F \longrightarrow$ chiral condensate enhancement

probes anomalous dimension γ_m needed by bilinear fermion masses $\sim \frac{1}{\Lambda^{2-\gamma_m}} \bar{q}q \bar{Q}Q$

(Future: Anomalous dimension $\gamma_{\mathcal{O}} = -\frac{d \log Z_{\mathcal{O}}(\mu)}{d \log \mu}$ for partial compositeness

from lattice non-perturbative renormalization of top-partner $\mathcal{O} \sim QQQ$)

- **Another example:** $\pi\pi$ scattering calculations on the lattice

\longrightarrow (complicated combinations of) electroweak LECs governing WW scattering

Evidence for dark matter

- Dark matter (DM) is definite physics beyond the standard model
Like composite Higgs, composite DM needs to explain known features of DM deduced from **consistent** (gravitational) evidence across all accessible scales
- Spiral galaxy rotation curves probe ‘small’ kiloparsec scales (1 pc \approx 3.26 ly)
Stars/gas far from center move faster than visible matter can explain
 \rightarrow local DM density $\sim 0.3 \text{ GeV/cm}^3$
- Similarly, galaxies within clusters move faster than expected, on Mpc scales
Lensing \rightarrow location of mass doesn’t match visible matter in clusters (e.g., bullet)
- Large-scale structure (Gpc scales) only reproduced by N -body simulations w/DM
- Cosmic microwave background (CMB) is largest visible distance scale $\approx 13.7 \text{ Gly}$
Power spectrum \rightarrow fluctuations at given angular scale (multipole),
third peak sensitive to dark (non-ionized) matter
- Fit CMB, large-scale structure, supernovae to six-parameter ‘ Λ CDM’ model
 $\rightarrow \sim 5\%$ ordinary matter, $\sim 25\%$ dark matter, $\sim 70\%$ dark energy
(most ordinary matter is interstellar H and He gas, most of remainder is stars)

Generic features of dark matter

- **Dark:** Electrically neutral, no direct coupling to photons
- **Matter:** In addition to gravitating, evolves like matter in expanding universe
Expansion \rightarrow metric $ds^2 = -dt^2 + a^2(t)dx_i^2$, scale factor $a(t)$ has $\frac{da}{dt} = aH(t) > 0$
(equivalently, cosmological redshift $z \equiv \frac{a_{\text{now}}}{a_{\text{past}}} - 1 > 0$)
Evolution \rightarrow Energy density $\rho \propto a^{-3(w+1)}$ where w is **equation of state**
relating pressure and energy density, $p = w\rho$
Matter: $w = 0 \rightarrow \rho \propto a^{-3} \sim \text{const./volume}$
Radiation (photons & **neutrinos**): $w = 1/3 \rightarrow \rho \propto a^{-4} \sim$ as above plus redshift
Cosmo. constant Λ : $w = -1 \rightarrow \rho = \text{const.}$
- **Stable:** Same ratio $\frac{\Omega_{DM}}{\Omega_B} \approx 5$ both today and at recombination
 \rightarrow little net decay or annihilation over ~ 13.7 billion years
- **Cold:** Large-scale structure formation from ‘bottom up’ \rightarrow **non-relativistic** DM
again rules out **neutrinos**, though very light axions are non-relativistic
- **Collisionless:** No (large) **dissipative** DM–DM (‘self-’)interactions,
would allow cooling \rightarrow **dark disks** that are **largely ruled out**
(Non-dissipative self-interactions would reduce substructure within galactic halos)
- Not clearly detected in ongoing **searches** \rightarrow at most weak DM–SM interactions

Motivations for composite dark matter

- Strong dynamics can produce stable massive particles (e.g., protons and nuclei)
- Dimensional transmutation \rightarrow natural hierarchy below Planck scale
- Possibility of joint solution to DM and EWSB from single new strong sector
(for now we will assume an elementary SM Higgs boson)
- Strong non-dissipative self-interactions may address [galactic structure issues](#)
(‘core-vs.-cusp’, ‘too-big-to-fail’, ‘missing satellites’, ...)
- Production of $\frac{\Omega_{DM}}{\Omega_B} \approx 5$ typically relies on non-gravitational DM–SM interactions,
while ongoing searches limit such interactions

Confinement of SM-charged ‘dark constituents’ into SM-neutral composite DM
could reconcile these two features

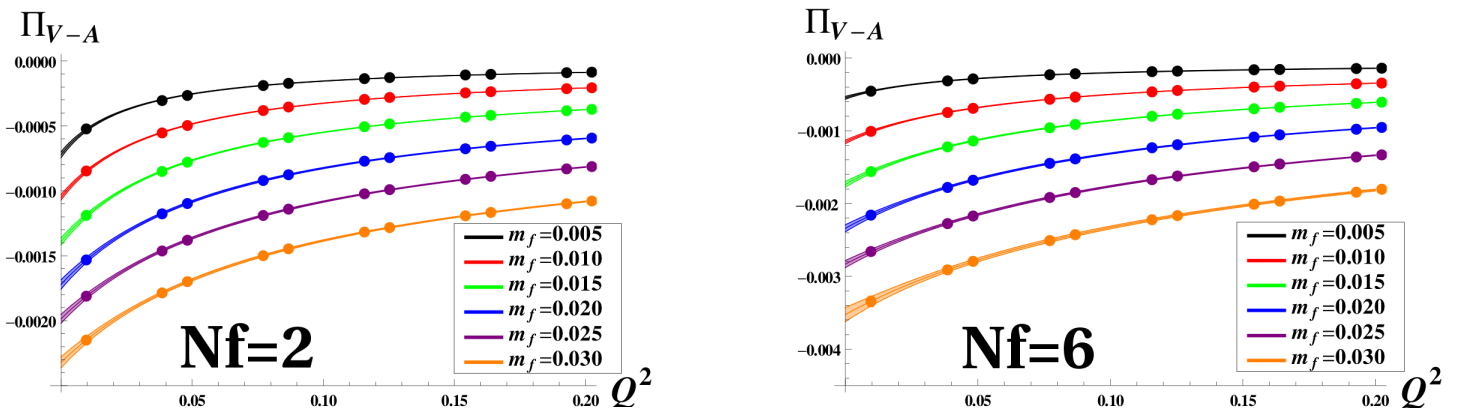


Figure 1: Fitting lattice data for the transverse vacuum polarization function $\Pi_{V-A}(Q^2)$ to Eq. 1, for SU(3) gauge theory with $N_F = 2$ (left) and 6 (right) domain wall fermions in the fundamental rep, from [arXiv:0910.2224](#). The smallest accessible euclidean momentum transfers are $Q^2 = \left(\frac{2\pi n}{N_t}\right)^2$ and $\left(\frac{2\pi n}{N_s}\right)^2$, where $N_t = 64$ and $N_s = 32$ are the temporal and spatial extents of the lattice, respectively. The different curves come from calculations with different fermion masses m_f , which need to be extrapolated to the chiral limit in which there are three exactly massless NGBs and $N_F^2 - 4$ massive PNGBs.

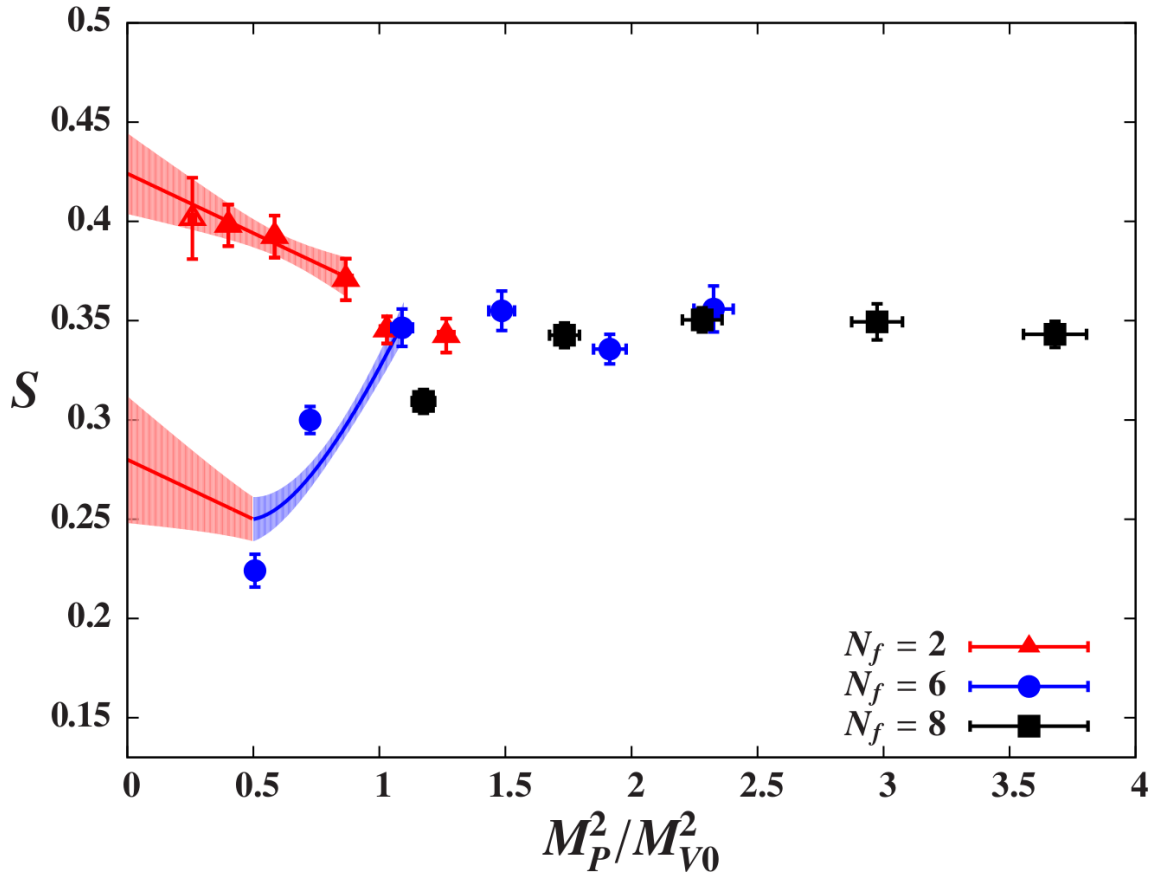


Figure 2: Results for the electroweak S parameter of SU(3) new strong dynamics with $N_F = 2, 6$ and 8 domain wall fermions in the fundamental rep, based on Fig. 11 of [arXiv:1405.4752](https://arxiv.org/abs/1405.4752). For $N_F = 8$ the Higgs could be either a dilaton or a PNGB, while $N_F < 7$ allows only the dilaton possibility with $f = v$. The $m \rightarrow 0$ extrapolated vector mass M_{V0} is used to define the UV cutoff $1/a$, which is finite but approximately matched between the three analyses. For sufficiently large $N_F \gtrsim 10$ the theory should flow to a chirally symmetric conformal fixed point in the IR, suggesting that V - A parity doubling may appear for intermediate $N_F = 6$ or 8 at sufficiently light PNGB masses $M_P \lesssim M_{V0}$, leading to a smaller S parameter as seen in the non-perturbative numerical results. However, the extra $N_F^2 - 4$ PNGBs lead S to diverge $\sim (N_F - 2) \log\left(\frac{M_{V0}^2}{M_P^2}\right)$ as $M_P \rightarrow 0$, so these must be kept massive, with only three exactly massless NGBs eaten by electroweak symmetry breaking. Here we fix by hand $M_P^2 \approx M_{V0}^2/2$ for the $N_F^2 - 4$ massive PNGBs, in order to produce the illustrative $N_F = 6$ extrapolation. The straightforward $N_F = 2$ extrapolation produces $S = 0.42(2)$ in agreement with scaling up QCD data. Unfortunately as N_F increases it becomes harder to access small M_P^2/M_{V0}^2 , further complicating the chiral extrapolations.

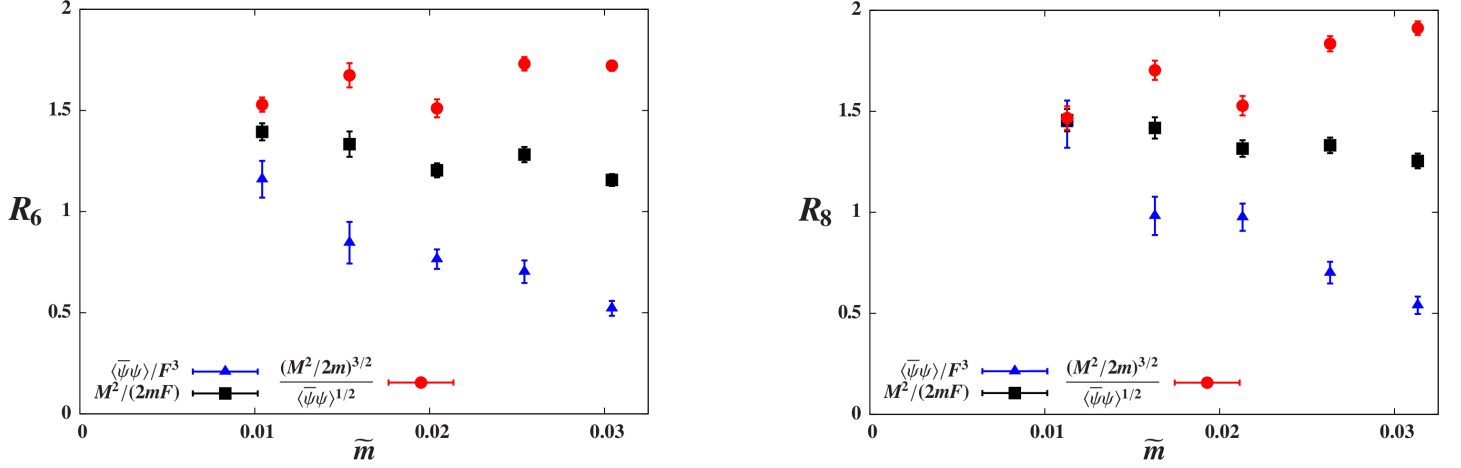


Figure 3: Enhancement of the chiral condensate in units of the decay constant $[\langle \bar{\psi}\psi \rangle / F^3]$ for SU(3) gauge theories with $N_F = 6$ (left) and 8 (right) domain wall fermions in the fundamental rep, from [arXiv:1405.4752](https://arxiv.org/abs/1405.4752). In each case we compare three quantities that reduce to $\langle \bar{\psi}\psi \rangle / F^3$ in the chiral limit $m \rightarrow 0$, each normalized by the corresponding $N_F = 2$ value computed with approximately matched M_{V0} standing in for the UV cutoff as above. The enhancement $R_F = \exp\left(2 \int_{\mu}^{\frac{d\mu}{\mu}} [\gamma_m^{(F)}(\mu) - \gamma_m^{(2)}(\mu)]\right) > 1$ implies a larger mass anomalous dimension $\gamma_m^{(F)}$ for F flavors compared to $N_F = 2$.

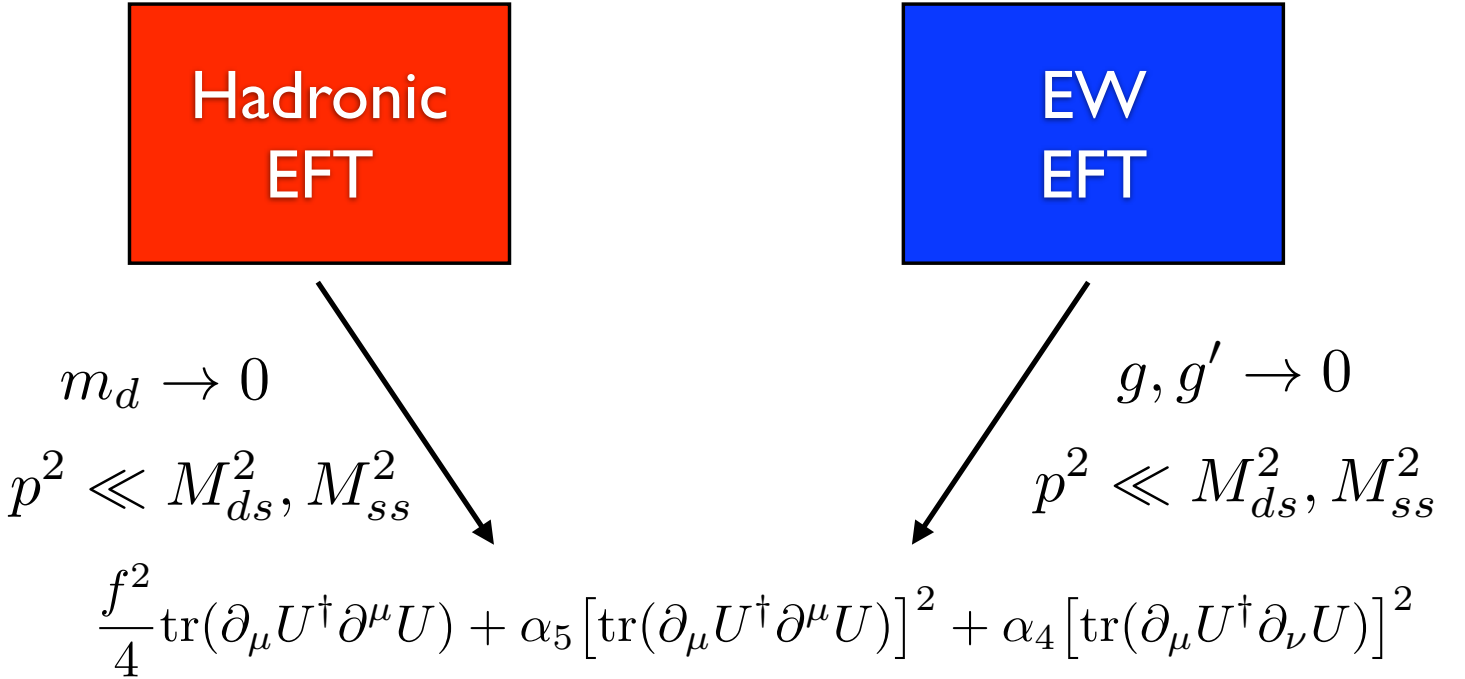


Figure 4: An illustration of relations between low-energy coefficients (LECs) of hadronic and electroweak EFTs (without including a light Higgs boson in the latter), focusing on the two electroweak LECs α_4 and α_5 that govern the scattering of like-sign WW bosons.

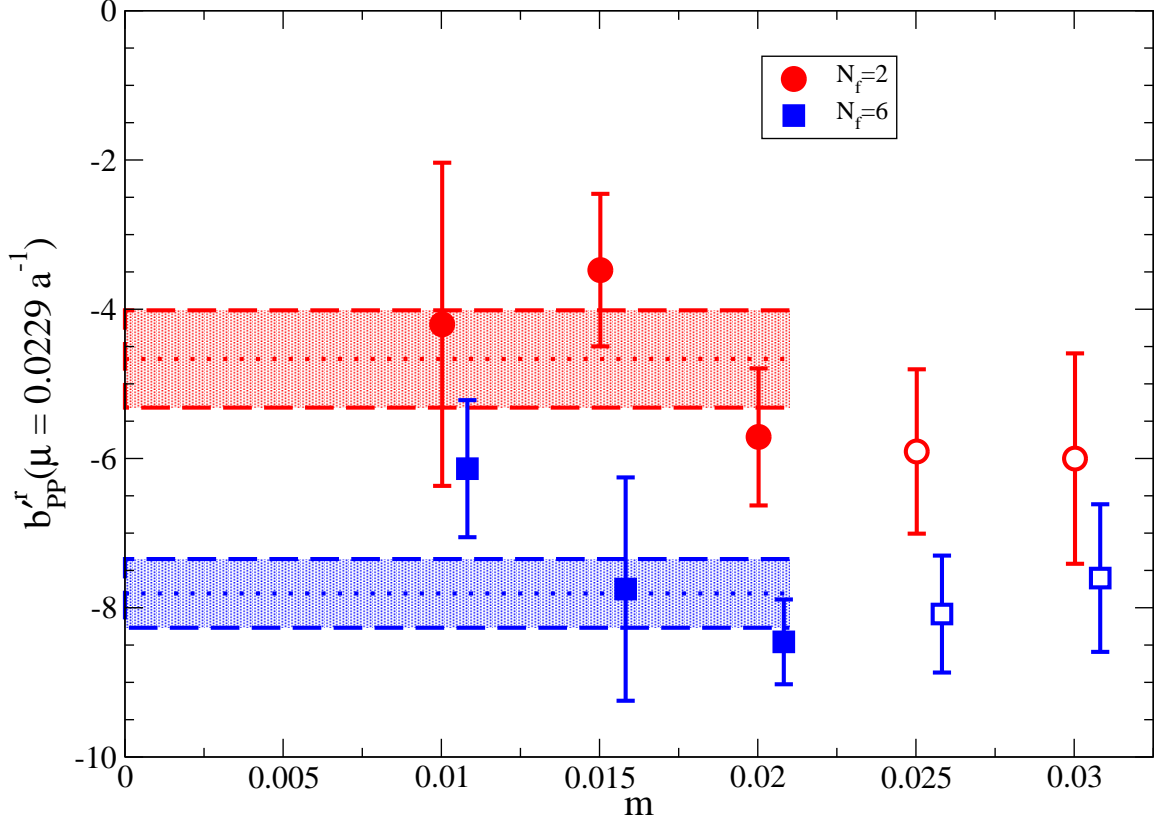


Figure 5: Results from [arXiv:1201.3977](https://arxiv.org/abs/1201.3977) for $b'_{PP} = -256\pi^2 [L_0 + 2L_1 + 2L_2 + L_3 - 2L_4 - L_5 + 2L_6 + L_8]$ (the linear combination of SU(3) chiral LECs L_i probed by maximal-isospin s-wave $\pi\pi$ scattering on the lattice), considering SU(3) gauge theories with $N_F = 2$ and 6 domain wall fermions in the fundamental rep. The electroweak LEC α_4 corresponds to the SU(2) chiral LEC ℓ_2 that depends on L_0 and L_2 ; similarly α_5 corresponds to ℓ_1 that depends on L_0 , L_1 and L_2 . Measurements of more processes (e.g., d-wave scattering or non-maximal isospin channels) would be needed to isolate these LECs. The more negative $N_F = 6$ results for b'_{PP} correspond to a slightly smaller $\pi\pi$ scattering length compared to $N_F = 2$.